Deep nets for making sense of ambient noise?

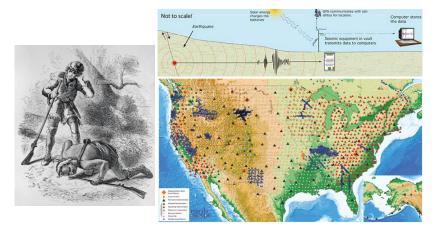
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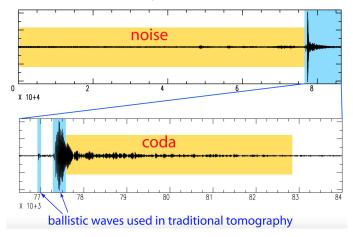
Example: ambient seismic noise



Credit: EarthScope, usarray.org

Noise, ballistic waves, and coda

one day of seismic record



Credit: Nikolai Shapiro, Michel Campillo

What information is present in ambient seismic noise?

Wave equation with random forcing

$$\left(\frac{1}{c^2(x)}(a+\partial_t)^2-\Delta\right)u(x,t)=n(x,t)$$

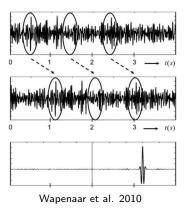
Space-time white noise: Gaussian with $\langle n(x,t)n(y,s)\rangle = \delta(t-s)\delta(x-y).$

- Pointwise: u(x, t) has Gaussian statistics in t
- Pairwise: u(x, t) and u(y, s) may be strongly correlated
 ⇒ may contain info about c(x)

Cross-correlations

Cross-correlations are robust to source randomness. Their peaks indicate traveltimes.

- Time reversal (phys): Fink et al. 1993.
- Time reversal (math) Bal, Papanicolaou, Ryzhik, 2002; Bal, Ryzhik 2003
- CINT imaging: Borcea, Papanicolaou, Tsogka, 2003, 2005
- Seismic interf. (phys): Weaver et al. 2001; Campillo, Paul 2003; Snieder, 2004; Wapenaar et al. 2006
- Seismic interf. (math): Bardos et al. 2008; Colin de Verdiere 2009; Garnier, Papanicolaou, Solna 2009+



Sample mathematical result (Bardos et al. 2008)

Assume a space-time white noise source. Let

$$C_T(\tau, x, y) = \frac{1}{T} \int_0^T u(x, t) u(y, t + \tau) dt.$$

Then

$$\partial_{\tau} \langle C_{T}(\tau, x, y) \rangle = -\frac{e^{-a|\tau|}}{4a} \operatorname{sgn}(\tau) G(|\tau|, x, y)$$

where

$$\left(\frac{1}{c^2(x)}\partial_t^2-\Delta\right)G(t,x,y)=\delta(t)\delta(x-y).$$

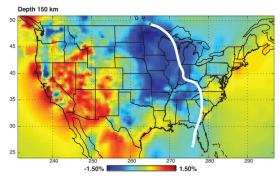
(Extensions: ergodicity, or large T limit, in the high freq regime.)

Seismic interferometry

In turn,

- G(t, x, y) is singular near t = traveltime(x, y)
- traveltime(x, y) is related to c(x)

Hence: determine c(x) by traveltime tomography



Credit: EarthScope, usarray.org

Open questions

Are c.c always the right idea?

What if the source distribution is **not isotropic**?

Is there physics that the lag of a cross-correlation peak doesn't see?

Can we avoid traveltime tomography to get c(x)?

What to do with combinations of 3 or more sensors?

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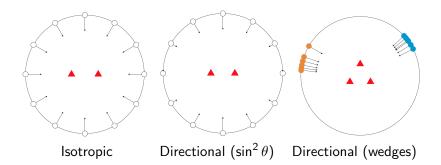
Deep nets provide "cheap" answers to most such questions (no guarantees, no physics)

Experimental setups

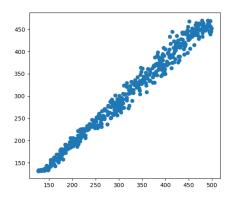
- Traveltime inversion with anisotropic sources, homogeneous media
- Identification of source distributions
- Inhomogeneous wavespeed inversion

Training: For homogeneous media, generate new data on the fly, seen once and never again. For inhomogeneous media, direct solve with conventional SGD + data augmentation

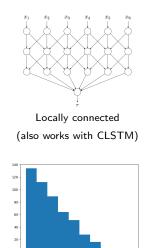
Traveltime inversion



Traveltime inversion: 2 sensors suffice



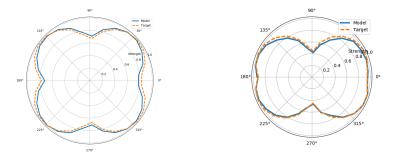
Estimated vs true traveltime (test samples)



Distribution of relative errors

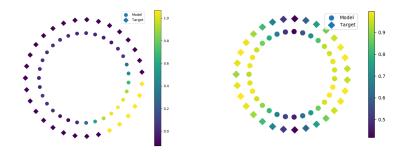
0.00 0.02 0.04 0.06 0.08 0.10 0.12 0.14

Inversion of source directionality



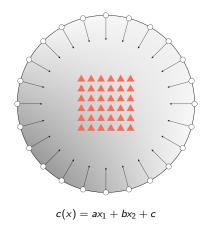
Estimated vs true source strength of the form $a \sin \theta + b \cos \theta$ (test samples). Need CLSTM architecture.

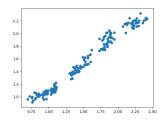
More Source Directionality



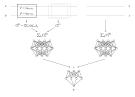
Left: network trained in a wedge distribution (with three sensors). Right: same as previous slide.

Inversion of heterogeneous wavespeeds



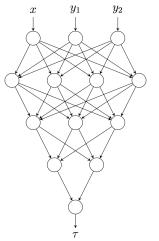


Estimated vs true traveltime

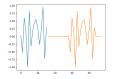


Relational network

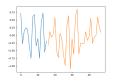
What are the nodes actually computing? (1)



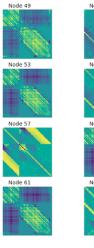
Noiseless:



Noisy:



What are the nodes actually computing? (2)



Node 50

Node 54

Node 58



Node 62



Node 63



Node 51

Node 55





Node 56



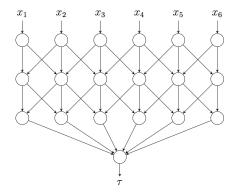
Node 60



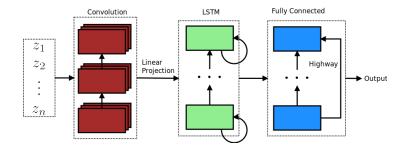
Node 64

Regress quadratic monomial on hidden nodes.

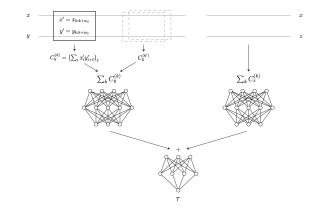
Architectures: Locally connected



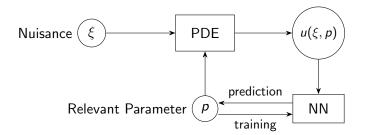
Architectures: CLSTM



Architectures: Relational



Automatic De-parameterization



- ξ : random source
- *p*: local wave speed, source distribution, etc
 u(ξ, *p*): passive seismic data

Goal: create representations of the data $u(\xi, p)$ that are mostly *invariant* to ξ and have enough information to recover p. In limited situations. this is what cross-correlations does.

Conclusions

• Blessing:

Data-centric predictions without a forward or inverse physical model (automatic de-parametrization of forward simulations that involve nuisance parameters that cannot be estimated during the inversion)

• Curse:

Physical content is (at best) buried in the network

• Curse:

Lack of performance guarantees

(possible overfitting, testing out of sample)

Outlook: characterization of scatterers, better performance in the inhomogeneous case

References I

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